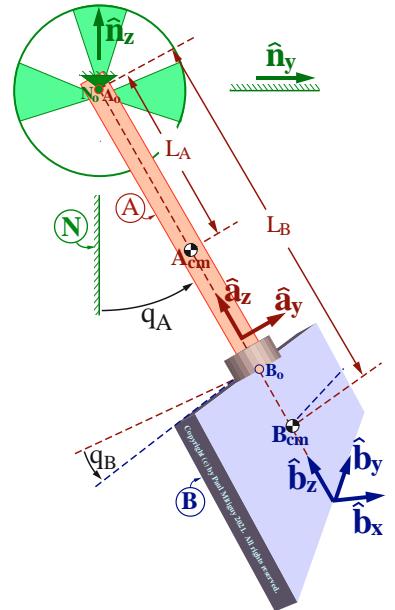


Motivating example (MIPSI): Babyboot

Model

Shown right is a model of a swinging babyboot (uniform plate B) attached by a shoelace (thin uniform rigid rod A) to a fixed rigid support N. Rod A is attached to N by a revolute joint at point N_o of N. B is attached to A with a 2nd revolute joint at point B_o so B can rotate freely about A's axis. Note: The revolute joints' axes are **perpendicular**, not parallel.

- **Bodies:** The rod and plate are **rigid** (inflexible/undeformable).
- **Connections:** Revolute joints are **ideal** (massless, frictionless, no slop).
- **Forces:** **Earth's gravity** is uniform and constant. Other contact forces (e.g., **air resistance** and solar/light pressure) and distance forces (e.g., electromagnetic and other gravitational) are negligible.
- **Newtonian reference frame:** Earth.



Identifiers

Right-handed sets of unit vectors $\hat{n}_x, \hat{n}_y, \hat{n}_z; \hat{a}_x, \hat{a}_y, \hat{a}_z; \hat{b}_x, \hat{b}_y, \hat{b}_z$ are fixed in N, A, B, respectively, with $\hat{n}_x = \hat{a}_x$ parallel to the revolute axis joining A to N, \hat{n}_z vertically-upward, $\hat{a}_z = \hat{b}_z$ parallel to the rod's long axis (and the revolute axis joining B to A), and \hat{b}_z perpendicular to plate B.

Quantity	Symbol	Type	Value
Earth's gravitational constant	g	Constant	9.81 m/s ²
Distance between N_o and A_{cm}	L_A	Constant	7.5 cm
Distance between N_o and B_{cm}	L_B	Constant	20 cm
Mass of A	m^A	Constant	0.01 kg
Mass of B	m^B	Constant	0.1 kg
A's moment of inertia about A_{cm} for \hat{a}_x	I^A	Constant	0.05 kg*cm ²
B's moment of inertia about B_{cm} for \hat{b}_x	I_x^B	Constant	2.5 kg*cm ²
B's moment of inertia about B_{cm} for \hat{b}_y	I_y^B	Constant	0.5 kg*cm ²
B's moment of inertia about B_{cm} for \hat{b}_z	I_z^B	Constant	2.0 kg*cm ²
Angle from \hat{n}_z to \hat{a}_z with $+\hat{n}_x$ sense	q_A	Dependent variable	Varies
Angle from \hat{a}_y to \hat{b}_y with $+\hat{a}_z$ sense	q_B	Dependent variable	Varies
Time	t	Independent variable	Varies

Note: Instructor worksheet at www.MotionGenesis.com ⇒ [Textbooks](#) ⇒ [Resources](#) completes the blanks above.

Physics

Physics from www.MotionGenesis.com ⇒ Get Started ⇒ Chaotic Pendulum (Babyboot).

The ODEs (ordinary differential equations) governing the motion of this mechanical system are⁴

$$\ddot{q}_A = \frac{2\dot{q}_A \dot{q}_B \sin(q_B) \cos(q_B) (I_x^B - I_y^B) - (m^A L_A + m^B L_B) g \sin(q_A)}{I^A + m^A L_A^2 + m^B L_B^2 + I_x^B \cos^2(q_B) + I_y^B \sin^2(q_B)}$$

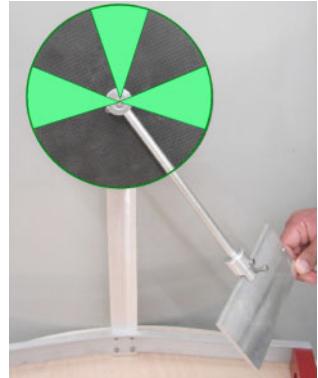
$$\ddot{q}_B = \frac{-\dot{q}_A^2 \sin(q_B) \cos(q_B) (I_x^B - I_y^B)}{I_z^B}$$

⁴Four methods to form equations of motion are: **Free-body diagrams** of A and B (which is inefficient as it introduces up to 10 unknown force/torque measures); D'Alembert's method (**MG road-maps** of Section 23.1) which efficiently forms the two equations shown for \ddot{q}_A and \ddot{q}_B (but requires a clever selection of systems, points, and unit vectors); **Lagrange's equations** (an energy-based method that automates D'Alembert's cleverness); **Kane's equations** (a modern efficient blend of D'Alembert and Lagrange).

Simplify and solve

The set of ODEs (ordinary differential equations) governing the babyboot's motion are coupled, nonlinear, 2nd-order ODEs. Computers have revolutionized the solution of ODEs and there are many numerical algorithms for solving ODEs (Euler's method, predictor-corrector, Runge-Kutta, etc). In addition, there are many programs (MATLAB®, MotionGenesis, WolframAlpha, etc.) that make it easy to solve ODEs.

```
% Computer solution of ODEs with MotionGenesis (with plotting).
%
% Variable qA'', qB'' % Angles and first/second time-derivatives.
%
qA'' = 2*( 508.89*sin(qA) - sin(qB)*cos(qB)*qA'*qB' ) / (-21.556 + sin(qB)^2)
qB'' = -sin(qB)*cos(qB)*qA'^2
%
Input tFinal = 10 sec, tStep = 0.02 sec, absError = 1.0E-07
Input qA = 90 deg, qB = 1.0 deg, qA' = 0.0 rad/sec, qB' = 0.0 rad/sec
Output t sec, qA degrees, qB degrees
%
ODE() solveBabybootODE
Plot solveBabybootODE.1[ 1, 2 ] % Plot qA (degrees) vs time t (seconds).
Plot solveBabybootODE.1[ 1, 3 ] % Plot qB (degrees) vs time t (seconds).
Quit
```



Interpret

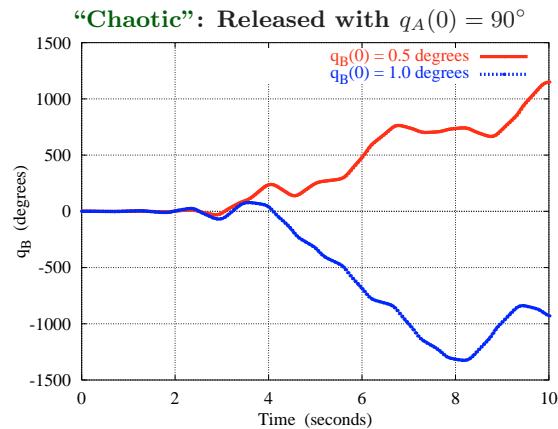
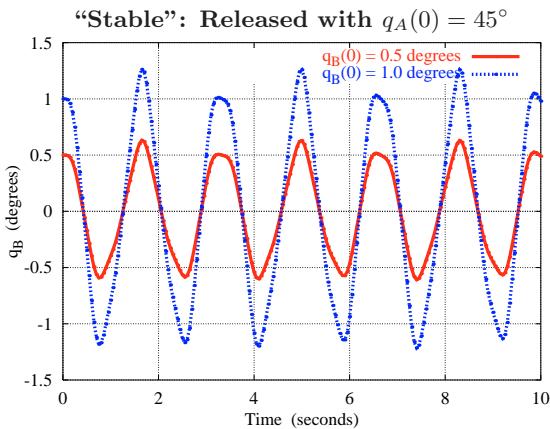
This simple system has strange non-intuitive motion. For certain initial values of q_A , plate B's motion is well-behaved and "stable" whereas for other initial values of q_A , B's motion is "**chaotic**" – meaning that a small variation in the initial value of q_B or numerical integration inaccuracies lead to dramatically different results (these ODEs test numerical integrators – the plots below required an integrator error of $\text{absError} = 1 \times 10^{-7}$).

The chart below and figure to the right shows this system's regions of stability (**black**) and instability (**green**).

Initial value of q_A	Stability	
$0^\circ \leq q_A(0) \leq 71.3^\circ$	Stable	black
$71.4^\circ \leq q_A(0) \leq 111.77^\circ$	Unstable	green
$111.78^\circ \leq q_A(0) \leq 159.9^\circ$	Stable	black
$160.0^\circ \leq q_A(0) \leq 180.0^\circ$	Unstable	green



The "**chaotic**" plot below shows q_B is **very** sensitive to initial values. A 0.5° change in the initial value $q_B(0)$ results in a 2000° difference in $q_B(t = 10)$!

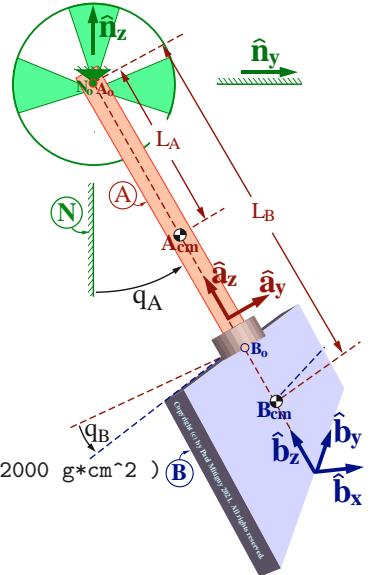


More information about this problem is in "Mechanical Demonstration of Mathematical Stability and Instability", *International Journal of Engineering Education (Journal of Mechanical Engineering Education)*, Vol. 2, No. 4, 1974, pp. 45-47, by Thomas Kane. Or visit www.MotionGenesis.com ⇒ [Get Started](#) ⇒ [Chaotic Pendulum \(Babyboot\)](#).

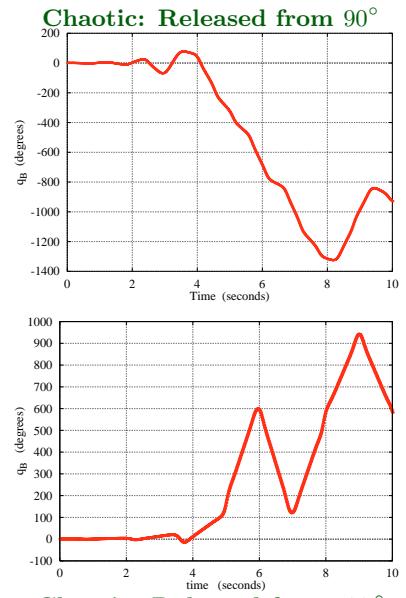
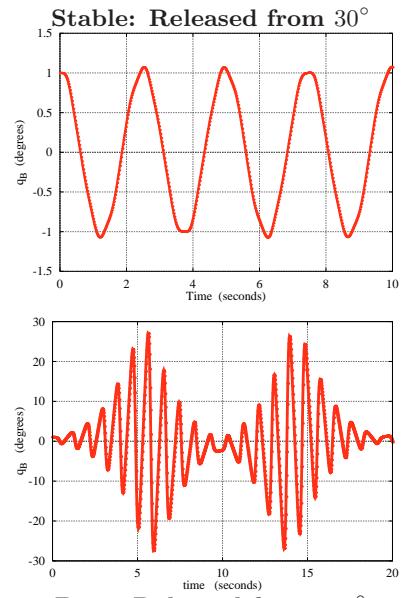
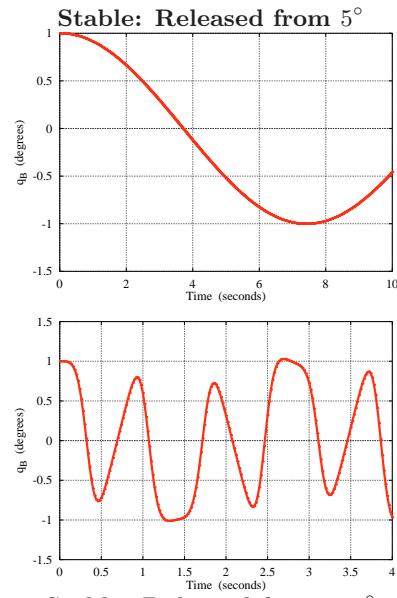
Physics

Physics from www.MotionGenesis.com ⇒ Get Started ⇒ Chaotic Pendulum (Babyboot).

```
% MotionGenesis file: MGBabybootDynamics.txt
% Problem: Analysis of 3D chaotic double pendulum.
% Copyright (c) 2009 Motion Genesis LLC. All rights reserved.
%-----
SetDigits( 5 ) % Number of digits displayed for numbers.
%-----
NewtonianFrame N % Earth.
RigidBody A % Upper rod.
RigidBody B % Lower plate.
%-----
Variable qA'' % Pendulum angle and its time-derivatives.
Variable qB'' % Plate angle and its time-derivative.
Constant LA = 7.5 cm % Distance from pivot to A's mass center.
Constant LB = 20 cm % Distance from pivot to B's mass center.
Constant g = 9.81 m/s^2 % Earth's gravitational acceleration.
A.SetMassInertia( mA = 10 grams, IAx = 50 g*cm^2, IAy, IAz )
B.SetMassInertia( mB = 100 grams, IBx = 2500 g*cm^2, IBy = 500 g*cm^2, IBz = 2000 g*cm^2 )
%-----
% Rotational kinematics.
A.RotateX( N, qA )
B.RotateZ( A, qB )
%-----
% Translational kinematics.
Acm.Translate( No, -LA*Az> )
Bcm.Translate( No, -LB*Az> )
%-----
% Add relevant contact/distance forces.
System.AddForceGravity( -g*Nz> )
%-----
% Equations of motion via free-body-diagrams (MG road-maps).
Dynamics[1] = Dot( Ax>, System(A,B).GetDynamics(No) )
Dynamics[2] = Dot( Bz>, B.GetDynamics(Bcm) )
%-----
% Kinetic and potential energy.
KE = System.GetKineticEnergy()
PE = System.GetForceGravityPotentialEnergy( -g*Nz>, No )
MechanicalEnergy = KE + PE
%-----
% Optional: Equations of motion with Kane's method.
SetGeneralizedSpeed( qA', qB' )
KaneDynamics = System.GetDynamicsKane()
isSameDynamics = IsSimplifyEqual( Dynamics, KaneDynamics )
%-----
% Optional: Equations of motion with Lagranges's method.
SetGeneralizedCoordinates( qA, qB )
LagrangeDynamics = System.GetDynamicsLagrange( SystemPotential = PE )
isSameDynamics := IsSimplifyEqual( Dynamics, LagrangeDynamics )
%-----
% Solve dynamics equations for qA'', qB''.
Solve( Dynamics = 0, qA'', qB'' )
%-----
% Integration parameters and initial values.
Input tFinal = 10 sec, tStep = 0.02 sec, absError = 1.0E-07, relError = 1.0E-07
Input qA = 90 deg, qA' = 0.0 rad/sec, qB = 1.0 deg, qB' = 0.0 rad/sec
%-----
% List output quantities and solve ODEs.
Output t sec, qA deg, qB deg, MechanicalEnergy Joules
ODE() MGBabybootDynamics
%-----
% Record input together with responses
Save MGBabybootDynamics.html
Quit
```



Investigation of stability: More simulation results



Stable: Released from 135°

Beat: Released from 158°

Chaotic: Released from 177°